DIFFSTOCHASTIC EQUIVALENT LINEARIZATION BASED ON THE FOKKER-PLANK EQUATION APPROACH

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1. Introduction

Stochastic equivalent linearization is the most popular approach to the approximate analysis of non-linear systems under random excitations. Over many years the original version of Gaussian equivalent linearization (GEL) has been developed by many authors, see e.g. [Atalik & Utku, 1976], [Casciati & Paraveilli, 1986], [Roberts & Spanos, 1990], [Zhang et al., 1991], [Anh & Schiehlen, 1995]. In order to improve the accuracy of GEL different techniques have been proposed, see e.g. [6, 7, 8].

In the paper a technique for determining the coefficients of the linearized equivalent equation based on the Fokker-Plank equation approach is presented. The investigation is then applied to Duffing and Vanderpol oscillations under a zero mean Gaussian white noise.

2. Equivalent linearization criterion

Consider a single-degree-of-freedom mechanical system, whose motion is described by the equation:

\[ \ddot{x} + 2h \dot{x} + \omega^2 x + \varepsilon f(x, \dot{x}) = \sigma \dot{\xi}(t), \]  

(2.1)

wherein the symbols have their customary meanings, \( f \) is a non-linear function of \( x \) and \( \dot{x} \), \( \omega \), \( h \), \( \sigma \) are positive constants, and \( \varepsilon \) is a positive parameter. The random excitation \( \xi(t) \) is a Gaussian white noise process of unit intensity

\[ E(\dot{\xi}(t), \dot{\xi}(t + \tau)) = \delta(\tau), \]  

(2.2)

where \( E(\cdot) \) denotes the expectation operation, and \( \delta(\tau) \) is Dirac-Delta function.

Following the linearization method, we introduce new linear terms in the expression of the equation (2.1)

\[ \ddot{x} + (2h + \varepsilon \mu) \dot{x} + (\omega^2 + \varepsilon \lambda)x + \varepsilon (f(x, \dot{x}) - \mu \dot{x} - \lambda x) = \sigma \dot{\xi}(t). \]  

(2.3)

The linearized equation takes the form

\[ \ddot{x} + (2h + \varepsilon \mu) \dot{x} + (\omega^2 + \varepsilon \lambda)x = \sigma \dot{\xi}(t). \]  

(2.4)

There are some criteria for determining the coefficients \( \mu, \lambda \), see e.g. [6, 7, 8]. In the paper an alternative approach to GEL is proposed as follows.
According to the classical approach of the averaging method the state coordinates \((x, \dot{x})\) are transformed to the pair of amplitude and phase \((a, \varphi)\) by the change

\[
x(t) = a \cos \varphi, \quad \dot{x}(t) = -a\omega \sin \varphi.
\]  

(2.5)

The Ito differential equations for \(a\) and \(\varphi\) are obtained from (2.3) [Mitropolskii et al., 1992]

\[
da = \frac{1}{\omega} \left[ (2h + \varepsilon\mu)\dot{x} + \varepsilon\lambda x \right] \sin \varphi
\]

\[+ \frac{\varepsilon}{\omega} \left( f - \lambda x - \mu \dot{x} \right) \sin \varphi + \frac{\sigma^2}{2\omega^2 a} \cos^2 \varphi \right] dt - \frac{a}{\omega} \sin \varphi d\xi(t),
\]

(2.6)

\[
d\varphi = \left[ \omega + \frac{1}{\omega a} \left[ ((2h + \varepsilon\mu)\dot{x} + \varepsilon\lambda x) \cos \varphi\right.ight.
\]

\[+ \varepsilon \left( f - \lambda x - \mu \dot{x} \right) \cos \varphi - \frac{\sigma^2}{2\omega^2 a^2} \sin^2 \varphi \right] dt - \frac{1}{\omega a} \cos \varphi d\xi(t).
\]

(2.6)

The Fokker-Planck equation (FP equation) for the stationary probability density function \(W(a, \varphi)\) corresponding to the system (2.6) takes the form:

\[
A_n[W] = \omega \frac{\partial W}{\partial \varphi} [K_1, K_2] W + \varepsilon \left\{ \frac{\partial}{\partial a} \left[ \frac{1}{\omega} \left( f - \lambda x - \mu \dot{x} \right) \sin \varphi \right] W \right\}
\]

\[+ \varphi \left[ \omega + \frac{1}{\omega a} \left[ ((2h + \varepsilon\mu)\dot{x} + \varepsilon\lambda x) \cos \varphi\right. \right.
\]

\[+ \varepsilon \left( f - \lambda x - \mu \dot{x} \right) \cos \varphi - \frac{\sigma^2}{2\omega^2 a^2} \sin^2 \varphi \right] dt - \frac{1}{\omega a} \cos \varphi d\xi(t).
\]

(2.7)

where \([K_1, K_2] W\) denotes the following linear differential operator

\[
[K_1, K_2] W = \frac{\partial}{\partial a} (K_1 W) + \frac{\partial}{\partial \varphi} (K_2 W)
\]

\[= - \frac{1}{2} \left[ \frac{\partial^2}{\partial a^2} (K_{11} W) + 2 \frac{\partial^2}{\partial a \partial \varphi} (K_{12} W) + \frac{\partial^2}{\partial \varphi^2} (K_{22} W) \right]
\]

(2.8)

and

\[
K_1 = \frac{\sigma^2}{4\omega^2 a} - \left( \hbar + \frac{\varepsilon\mu}{2} \right) a + \left[ \frac{\sigma^2}{4\omega^2 a} + \left( \hbar + \frac{\varepsilon\mu}{2} \right) a \right] \cos 2\varphi + \frac{\varepsilon\lambda}{2\omega} a \sin 2\varphi,
\]

\[
K_2 = \frac{\varepsilon\lambda}{2\omega} + \frac{\varepsilon\lambda}{2\omega} \cos 2\varphi \left( \hbar + \frac{\varepsilon\mu}{2} + \frac{\sigma^2}{2\omega^2 a^2} \right) \sin 2\varphi,
\]

\[
K_{11} = \frac{\sigma^2}{2\omega^2 a^2} \left( 1 - \cos 2\varphi \right); \quad K_{12} = \frac{\sigma^2}{2\omega^2 a^2} \sin 2\varphi; \quad K_{22} = \frac{\sigma^2}{2\omega^2 a^2} \left( 1 + \cos 2\varphi \right).
\]

(2.9)

Let \(W_L(a, \lambda, \mu)\) be a solution of the FP equation corresponding to the linearized systems (2.4). If \(A_n[W_L(a, \lambda, \mu)] = 0\), then \(W_L(a, \lambda, \mu)\) is also a solution of the FP equation (2.7) which corresponds to the non-linear systems (2.3). However, it is seen from (2.7) that in general \(A_n[W_L(a, \lambda, \mu)]\) differs from zero. So, we minimise the error in mean square as follows

\[
\int_0^{2\pi} \int_0^{2\pi} \left\{ A_n[W_L(a, \lambda, \mu)] \right\}^2 \, da \, d\varphi \rightarrow \min.
\]

(2.10)

The FP equation corresponding to the linearized equation (2.4) gives the solution

\[
W_L(a, \lambda, \mu) = ca \exp \left\{ - \frac{(\sigma^2 + \varepsilon\lambda)(2h + \varepsilon\mu) a^2}{\sigma^2 a^2} \right\},
\]

(2.11)
where $c$ is a normalization constant.

Substituting (2.11) into (2.10) one gets

$$H(\lambda, \mu) = \int_{0}^{2\pi} \int_{0}^{2\pi} \left\{ \frac{\partial}{\partial \lambda} \left[ \frac{1}{\omega} (f - \lambda x - \mu \dot{x}) \sin \varphi W_{1} \right] + \frac{\partial}{\partial \varphi} \left[ \frac{1}{\omega} (f - \lambda x - \mu \dot{x}) \cos \varphi W_{1} \right] \right\} \, da \, d\varphi \rightarrow \min.$$  \hfill (2.12)

The condition (2.12) gives a criterion for determining the linearization coefficients $\lambda$ and $\mu$.

3. Application

In order to elucidate this approach we consider two oscillators with non-linear stiffness and damping, under white noise excitation.

3.1. Duffing oscillator

Consider the following equation

$$\ddot{x} + 2h\dot{x} + \omega^2 x + \epsilon \gamma x^3 = \sigma \dot{\xi}(t),$$  \hfill (3.1)

where $\gamma = \text{const}$. In this case one gets

$$f(x, \dot{x}) = \gamma x^3,$$  \hfill (3.2)

and the linearized equation is

$$\ddot{x} + 2h\dot{x} + (\omega^2 + \epsilon \lambda)x = \sigma \dot{\xi}(t).$$  \hfill (3.3)

So, one gets probability density function (2.11) in the form

$$W_{L}(a, \lambda) = c a \exp \left\{ - \frac{(\omega^2 + \epsilon \lambda)^2 h}{\sigma^2} a^2 \right\}. \hfill (3.4)$$

Substituting (3.2), (3.4) into (2.12) and using (2.5), after some calculations one gets:

$$H(\alpha) = \pi \sigma^2 \left( \frac{5\gamma^2 \alpha^2}{16 \omega^2} J_{10} + \frac{\gamma \lambda \alpha^2}{\omega^2} J_{8} + \frac{\lambda^2 \alpha^2}{\omega^2} J_{6} \right), \hfill (3.5)$$

where

$$\alpha = \frac{2h(\omega^2 + \epsilon \lambda)}{\sigma^2}. \hfill (3.6)$$

The expressions $J_{i}$ \((i = 0, 2, 4, \ldots)\) are shown in the Appendix 1.

Substituting $J_{6}$, $J_{8}$, $J_{10}$ and (3.6) into (3.5) one gets, finally

$$H(\alpha) = \pi \sqrt{2\pi} \, \sigma^2 \left\{ \frac{4725 \gamma^2}{2^{10} \omega^2} \, \frac{1}{\alpha^2 \sqrt{\alpha}} - \frac{105}{2^{10} \epsilon} \, \frac{1}{\alpha^2 \sqrt{\alpha}} \right\} \right. \hfill (3.7)

+ \left( \frac{105}{2^{10} \epsilon} \, \frac{\sigma^2}{h} + \frac{15 \omega^2}{2^8 \epsilon^2} \right) \frac{1}{\alpha \sqrt{\alpha}} - \frac{15}{2^8 \epsilon} \left( \frac{\sigma^2}{h} \right)^2 \frac{1}{\alpha} + \frac{15}{2^{10} \epsilon^2 \sigma^2} \left( \frac{\sigma^2}{h} \right)^2 \sqrt{\alpha} \right\}.$$
In order to determine the smallest value of the function $H(\alpha)$ in the interval $\alpha \in (0, +\infty)$ we consider the following equation

$$
\frac{dH(\alpha)}{d\alpha} = \pi \sqrt{\pi} \left\{ - \frac{33075 \tau^2}{21^2 \omega^2} \frac{1}{\alpha^4 \sqrt{\alpha}} + \frac{525}{21 \omega} \frac{1}{\alpha^3 \sqrt{\alpha}} - \left( \frac{315}{21^2 \omega^2} \frac{\sigma^2}{\epsilon} \right) \frac{1}{\alpha \sqrt{\alpha}} + \frac{15}{2^2 \omega^2} \left( \frac{\sigma^2}{\alpha} \right) \frac{1}{\alpha^2 \sqrt{\alpha}} \right\} = 0.
$$

(3.8)

From the equation (3.8) one obtains the solution $\alpha = \alpha_{\text{min}} > 0$ at which the function (3.7) has the smallest value $H_{\text{min}} = H(\alpha_{\text{min}})$. The value $\alpha_{\text{min}}$ depending one $\omega$, $\gamma$, $\sigma^2$, $h$, $\epsilon$ can be determined numerically.

The mean square of displacement corresponding to $\alpha_{\text{min}}$ takes the form

$$
\langle x^2 \rangle_T = \int_0^{2\pi} \int_0^\infty x^2 W_L(a, \alpha_{\text{min}}) da dp = \frac{1}{2 \alpha_{\text{min}}}.
$$

(3.9)

The result obtained by the expression (3.9) $\langle x^2 \rangle_T$ is compared in the Tab.1 with the exact solution for values $\omega = 1$, $\gamma = 1$, $\sigma^2 = 4h$, and different values of $\epsilon$. In addition, the result obtained by the classical GEL $\langle x^2 \rangle_g$ is also shown. Obviously, the solution $\langle x^2 \rangle_T$ is much closer to the exact solution $\langle x^2 \rangle_e$ than the solution $\langle x^2 \rangle_g$.

<table>
<thead>
<tr>
<th>$N$</th>
<th>$\epsilon$</th>
<th>$\langle x^2 \rangle_e$</th>
<th>$\langle x^2 \rangle_g$</th>
<th>$\langle x^2 \rangle_T$</th>
</tr>
</thead>
<tbody>
<tr>
<td>1</td>
<td>0.1</td>
<td>0.8176</td>
<td>0.8054 (1.49%)</td>
<td>0.8289 (1.38%)</td>
</tr>
<tr>
<td>2</td>
<td>1.0</td>
<td>0.4679</td>
<td>0.4343 (7.19%)</td>
<td>0.4492 (3.99%)</td>
</tr>
<tr>
<td>3</td>
<td>10.0</td>
<td>0.1889</td>
<td>0.1667 (11.8%)</td>
<td>0.1745 (7.57%)</td>
</tr>
<tr>
<td>4</td>
<td>100.0</td>
<td>0.0650</td>
<td>0.0561 (13.6%)</td>
<td>0.0588 (9.54%)</td>
</tr>
</tbody>
</table>

3.2. Van der Pol oscillator

As a second example we consider a Vanderpol oscillator

$$
\ddot{x} + \omega^2 x + \epsilon(\gamma x^2 - 1)\dot{x} = \sigma \xi(t).
$$

(3.10)

In this case one gets

$$
f(x, \dot{x}) = (\gamma x^2 - 1)\dot{x}, \quad \gamma = \text{const},
$$

(3.11)

and the linearized equation takes the form

$$
\ddot{x} + \epsilon \mu \dot{x} + \omega^2 x = \sigma \xi(t).
$$

(3.12)

So, one gets the following probability density function

$$
W_L(a, \mu) = c \exp \left\{- \frac{\omega^2 \epsilon \mu a^2}{\sigma^2} \right\}.
$$

(3.13)

Substituting (3.11), (3.13) into (2.12) and noting (2.5), after some calculations one obtains

$$
H(\beta) = \pi \sqrt{2 \pi} \epsilon \sigma \left\{ \frac{29 \sigma^4}{2^8 \omega^2 \mu^4} \sqrt{\beta} + \frac{29 \sigma^2}{2^7 \epsilon \omega^2} \frac{1}{\beta^{\frac{3}{2}}} \right\}
$$

$$
+ \left( \frac{29}{2^8} + \frac{249 \gamma \sigma^2}{21^2 \epsilon \omega^2} \right) \frac{1}{\beta^{\frac{1}{2}}} \cdot \frac{1}{\beta^{\frac{3}{2}}} + \frac{2497 \gamma^2}{21^10} \cdot \frac{1}{\beta^{\frac{1}{2}}} + \frac{9375 \gamma^2}{21^16} \cdot \frac{1}{\beta^{\frac{3}{2}}} \right\}.
$$

(3.14)
where
\[ \beta = \frac{\sigma^2 \mu}{\sigma^2}. \]  

(3.15)

Differentiating the function (3.14) with respect to \( \beta \) one gets the following equation
\[
\frac{dH(\beta)}{d\beta} = \pi \sqrt{\pi} \sigma^2 \left[ \frac{29\sigma^4}{2^9 e^2 \omega^4} \cdot \frac{1}{\sqrt{\beta}} - \frac{29\sigma^2}{2^7 e \omega^2} \cdot \frac{1}{\beta \sqrt{\beta}} - \frac{1}{\beta^2 e^2 \omega^2} \left( \frac{87}{2^7} + \frac{747 \gamma \sigma^2}{2^{11} e^2 \omega^2} \right) \cdot \frac{1}{\beta^3 \sqrt{\beta}} \right] = 0.
\]  

(3.16)

The equation (3.16) gives the solution \( \beta = \beta_{\text{min}} > 0 \) at which the function (3.14) has the smallest value \( H_{\text{min}} = H(\beta_{\text{min}}) \) in the interval \( \beta \in (0, +\infty) \). The result obtained from the expression similar to the one (3.9) \( \langle \xi^2 \rangle_T \) is compared in the Tab. 2 with the simulation solutions \( \langle \xi^2 \rangle_{MC} \) taken in [Roy and Spanos, 1991] where \( \omega = 1, \gamma = 10, \varepsilon = 0.2 \) and for different values of \( \sigma^2 \). In addition, the results obtained by the classical GEL and the classical SAM (Stochastic averaging method) \( \langle \xi^2 \rangle_{SAM} \) [Roy & Spanos, 1991] are also shown. It is seen from the Table 2 that the solution \( \langle \xi^2 \rangle_T \) is much closer to the simulation \( \langle \xi^2 \rangle_{MC} \) than the solution \( \langle \xi^2 \rangle_{sg} \).

<table>
<thead>
<tr>
<th>( N )</th>
<th>( \sigma^2 )</th>
<th>( \langle \xi^2 \rangle_{MC} )</th>
<th>( \langle \xi^2 \rangle_{sg} )</th>
<th>( \langle \xi^2 \rangle_{SAM} )</th>
<th>( \langle \xi^2 \rangle_T )</th>
</tr>
</thead>
<tbody>
<tr>
<td>1</td>
<td>0.02</td>
<td>0.2080</td>
<td>0.137 (34%)</td>
<td>0.2055 (1.2%)</td>
<td>0.1859 (10.62%)</td>
</tr>
<tr>
<td>2</td>
<td>0.2</td>
<td>0.3600</td>
<td>0.279 (22%)</td>
<td>0.3402 (5.5%)</td>
<td>0.3445 (4.49%)</td>
</tr>
<tr>
<td>3</td>
<td>1.0</td>
<td>0.7325</td>
<td>0.552 (25%)</td>
<td>0.6433 (12.1%)</td>
<td>0.6903 (9.17%)</td>
</tr>
<tr>
<td>4</td>
<td>2.0</td>
<td>1.0310</td>
<td>0.759 (26%)</td>
<td>0.8750 (15.1%)</td>
<td>0.9152 (11.23%)</td>
</tr>
<tr>
<td>5</td>
<td>4.0</td>
<td>1.4540</td>
<td>1.051 (28%)</td>
<td>1.2040 (17.2%)</td>
<td>1.2916 (11.17%)</td>
</tr>
</tbody>
</table>

4. Conclusions

The problem for discovering new techniques to overcome the limitation of the classical GEL for both weak and strong non-linearity is of great interest. In the paper a technique to treat stationary response of non-linear systems under a zero mean Gaussian white noise is presented based on the Fokker-Plank equation approach. The proposed technique is then applied to Duffing and Van der Pol oscillators to show significant improvements over the accuracy of classical GEL.

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Appendix

1. Let
\[
J_{2n} = \int_{0}^{\infty} \alpha^{2n} e^{-2\alpha^2} d\alpha; \quad (\alpha > 0, \ n = 0, 1, 2, \ldots), \tag{A.1}
\]

one gets
\[
J_0 = \frac{\sqrt{2\pi}}{4\alpha}, \tag{A.2}
\]

and
\[
J_{2n+2} = \frac{2n + 1}{4\alpha} J_{2n} \quad (n = 0, 1, 2, \ldots). \tag{A.3}
\]
2. Property

Let the function \( y = f(x) \) be defined, continuously in the interval \( x \in (0, +\infty) \) and satisfy the following conditions

\[
\lim_{x \to 0} f(x) = +\infty; \quad \lim_{x \to +\infty} f(x) = +\infty.
\]

It is known that the function \( f(x) \) has the smallest value in the interval \( (0, +\infty) \). In particular, the functions \( H(\alpha) \) in (19) and \( H(\beta) \) in (26) have the property mentioned above, so both of them have the smallest values in the interval \( (0, +\infty) \).

References


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